



## **Air Force Maui Optical & Supercomputing Site (Technology Overview)**

**Dr. Keith Knox**

**Boeing LTS, Chief Scientist**

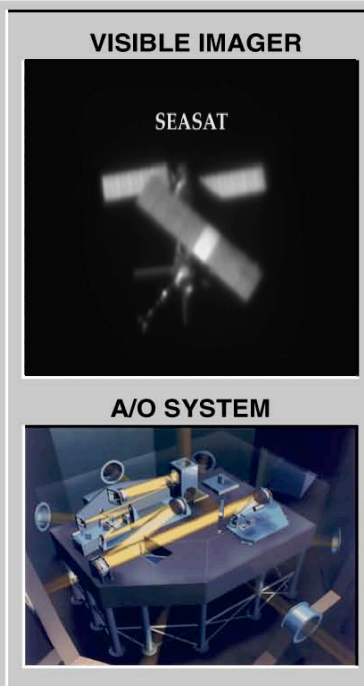
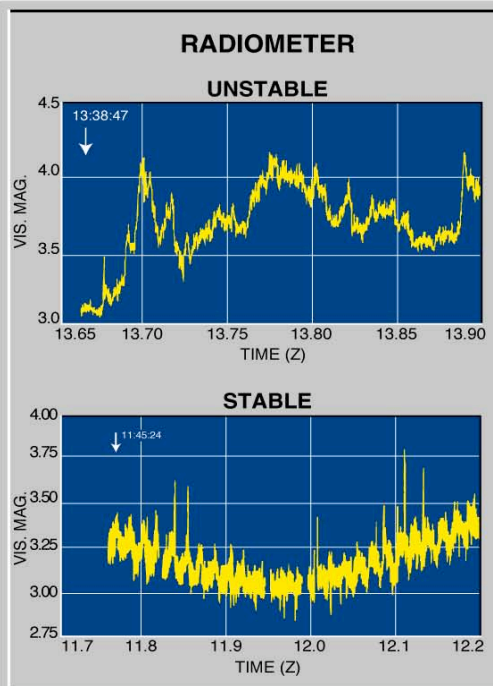
**Air Force Maui Optical & Supercomputing Site**



# 3.6-meter Telescope Sensor Suite Covers Visible through LWIR Spectrum



Data  
shows  
satellite  
orientation  
stability

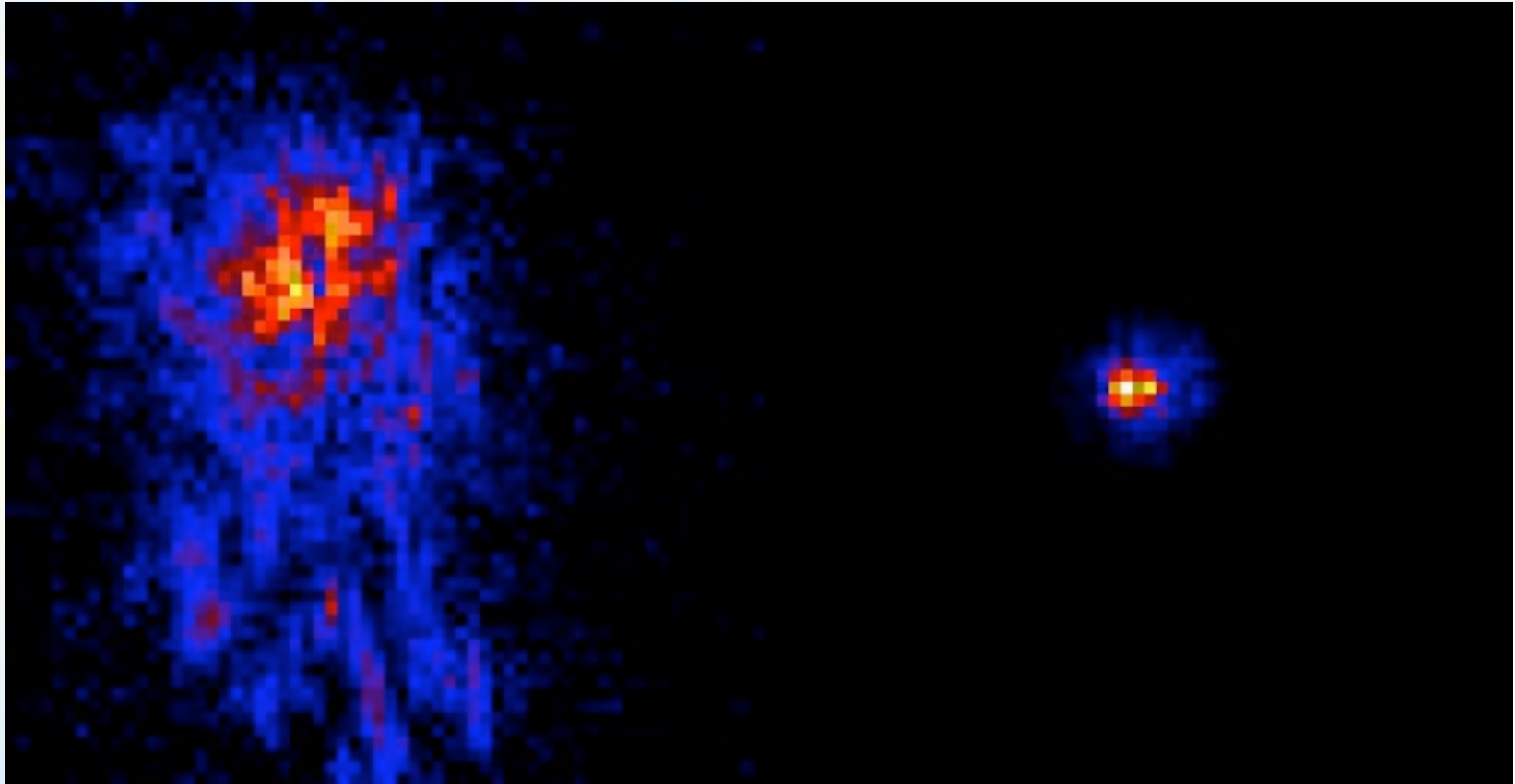


Field data  
results -  
using adaptive  
optics (A/O)  
compensation

Sensor	Aperture (cm)	Spectral Range (um)	Full FOV	Array Size	Detector Type
Wide FOV Acquisition	20	0.4 - 0.9	2.7 deg	512 <sup>2</sup>	ICCD
Narrow FOV Acquisition	58	0.4 - 0.9	0.45/0.125 deg	512 <sup>2</sup>	ICCD Visible: Si
MW Acquisition	45	3 - 5	0.3 deg		MWIR: InSb
Photometer / Radiometer	363	0.4 - 14	15 - 150 urad (Electronic)	128 <sup>2</sup>	LWIR: Si:As LWIR: Si:As
LWIR Imager	363	8 - 14	140 urad	200 <sup>2</sup>	Si:As
Visible Imager	363	0.4 - 0.9	50 - 120 - 200		



# Adaptive Optics: Real-time Correction of Atmospheric Distortions



Atmospheric distortions

Adaptive optics compensation



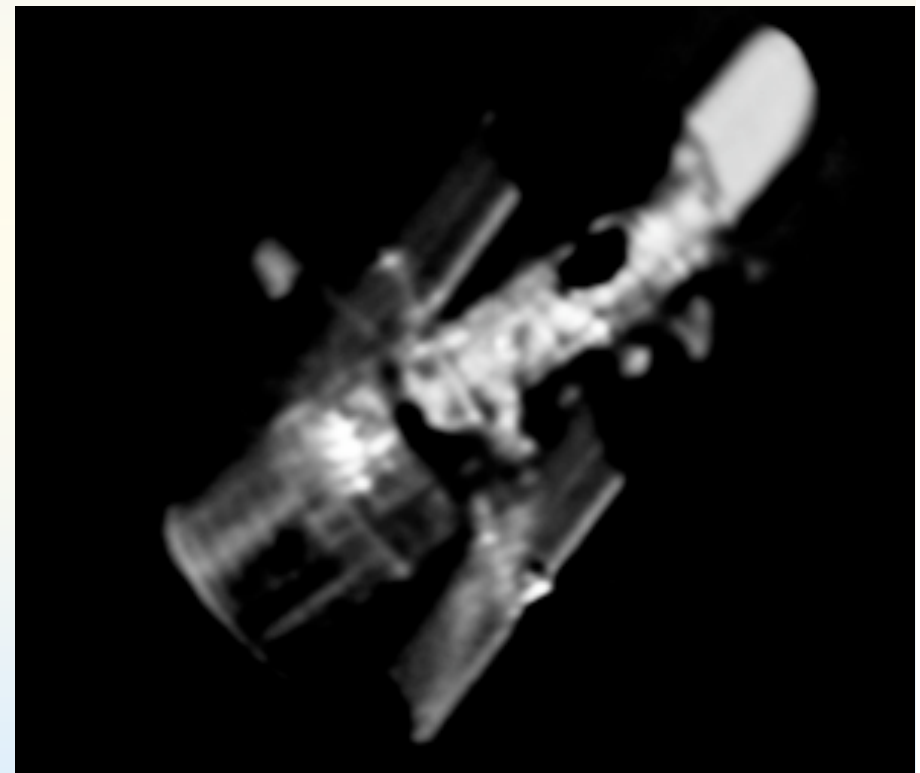
# A/O Adaptive Optics Imaging



**Provides high-resolution visible imagery**

Hubble Space Telescope

- **Adaptive Optics Compensation**
  - **Wavefront sensor**
  - **Deformable mirror**
- **Real-time correction**
  - **Atmosphere measured 200 cycles/second**
  - **Image recorded 5 frames/second**



**Adaptive Optics plus Multi-Frame  
Blind Deconvolution Processing**



# 3.6-meter Telescope LWIR Imager (Long Wave Infrared)



## Provides high-resolution thermal

- Produces Temperature Estimates of Objects in Orbit
- Dual Si:As Focal Plane Arrays  
8.1-9.1  $\mu\text{m}$  & 10.1-12.9  $\mu\text{m}$
- Background-Limited Sensitivity
- Virtually Diffraction-Limited
- Operates Simultaneously w/ Visible Imager

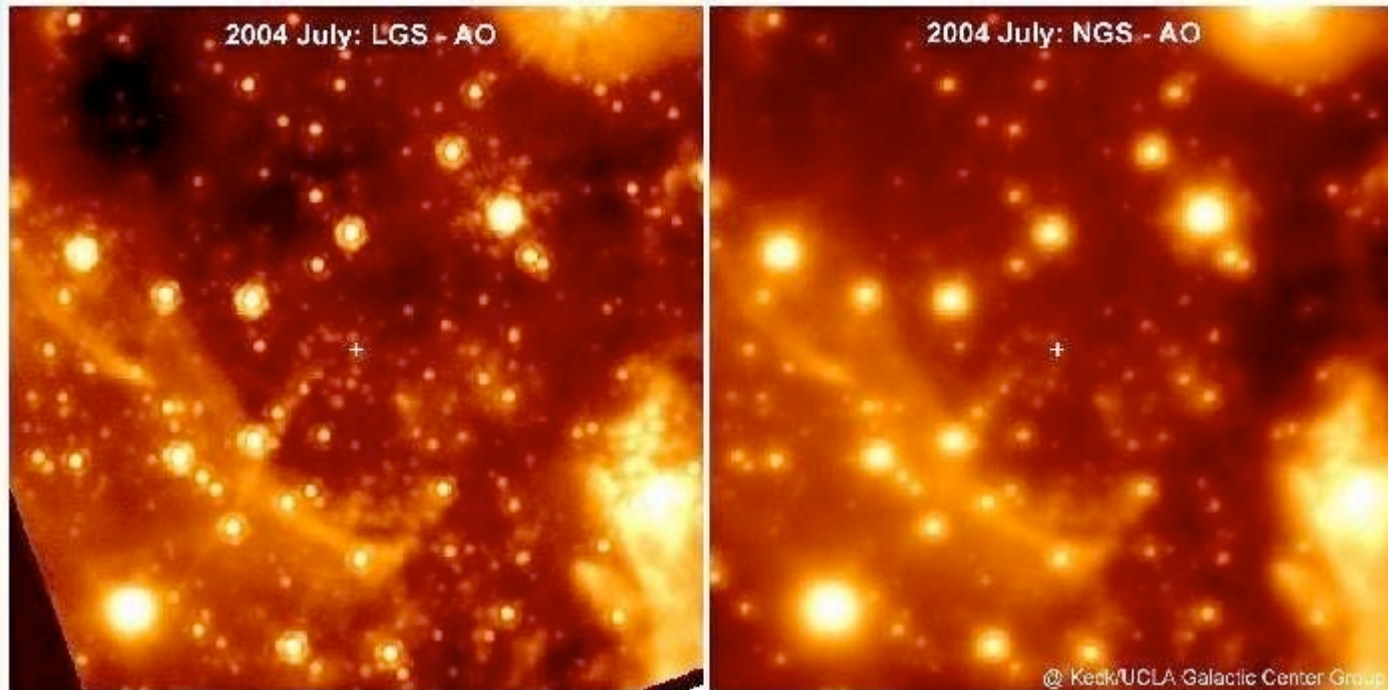


Developed by Raytheon,  
El Segundo





# Laser Guide Star Adaptive Optics



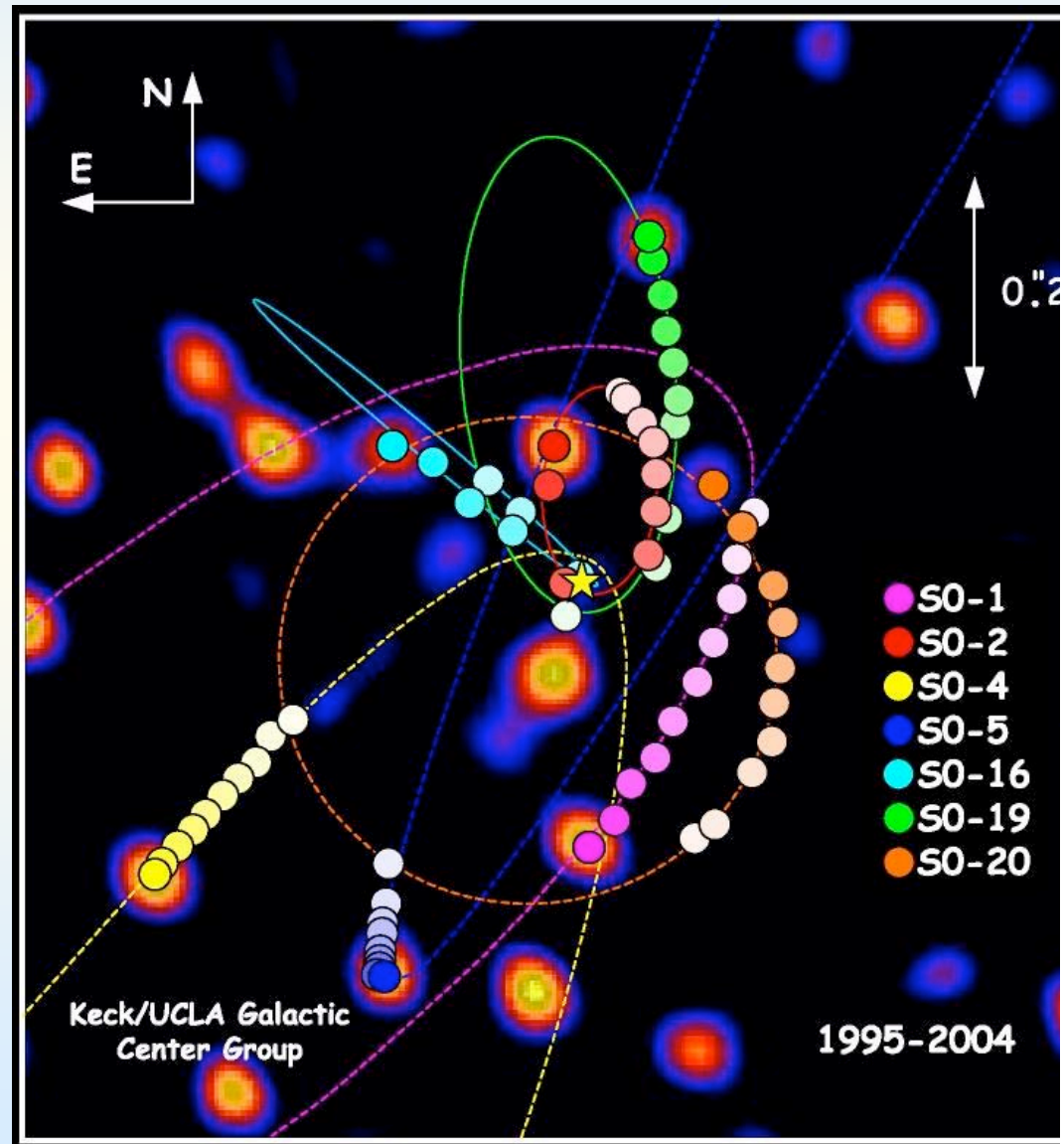
- **UCLA PRESS RELEASE (December 12th, 2005)** UCLA astronomers and colleagues published the first high-resolution images of the center of our Milky Way galaxy, including the area surrounding the supermassive black hole, using a new technology at the W.M. Keck Observatory in Hawaii.
  - “We have worked for years on techniques for ‘beating the distortions in the atmosphere’ and producing high-resolution images,” Andrea Ghez, UCLA professor of physics and astronomy said. “We are pleased to report the first Laser Guide Star Adaptive Optics observations of the center of our galaxy.”



# Discovery of Black Hole at Center of Milky Way Galaxy



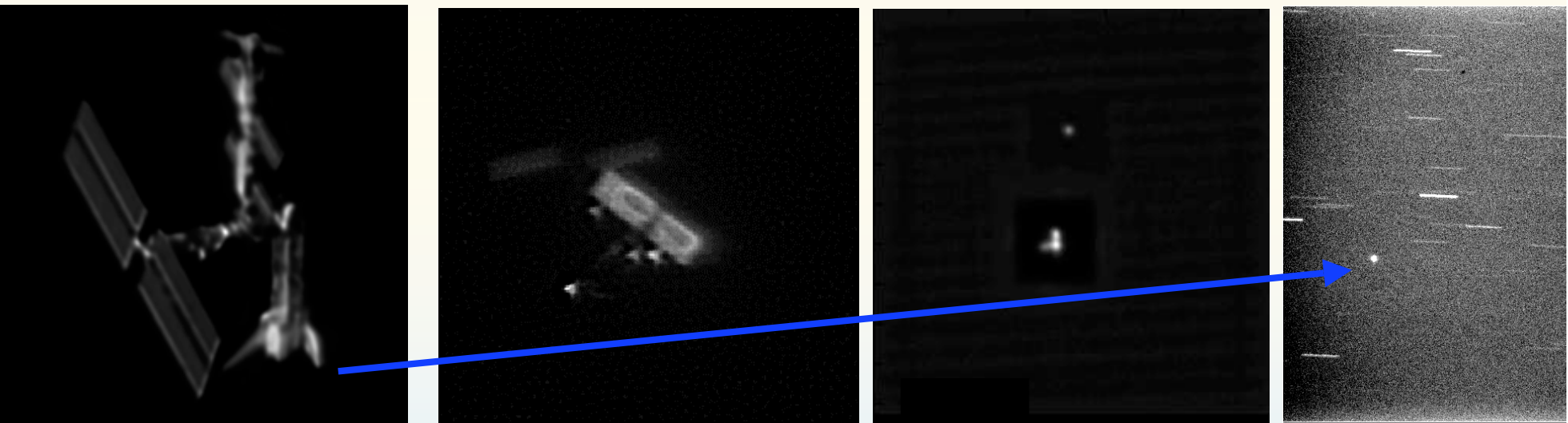
**MAUNA KEA, Hawaii  
(September 21st, 2000) In  
an historic first,  
astronomers from the  
University of California-  
Los Angeles have  
observed acceleration in  
the velocities of  
individual stars orbiting  
the gigantic black hole at  
the center of our galaxy.  
The astronomers are led  
by Andrea Ghez, UCLA  
professor of physics and  
astronomy.**





# Satellite Characterization

- As angular size of satellite decreases
  - Smaller satellite
  - Greater distance



- As telescope aperture decreases
- Resolution decreases, until the satellite is completely unreso





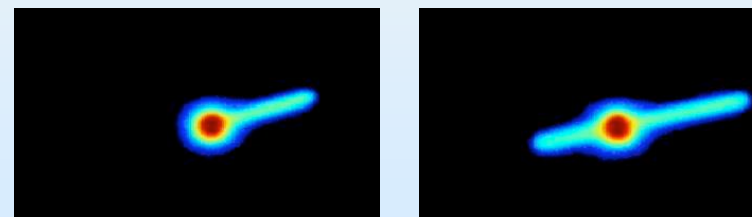
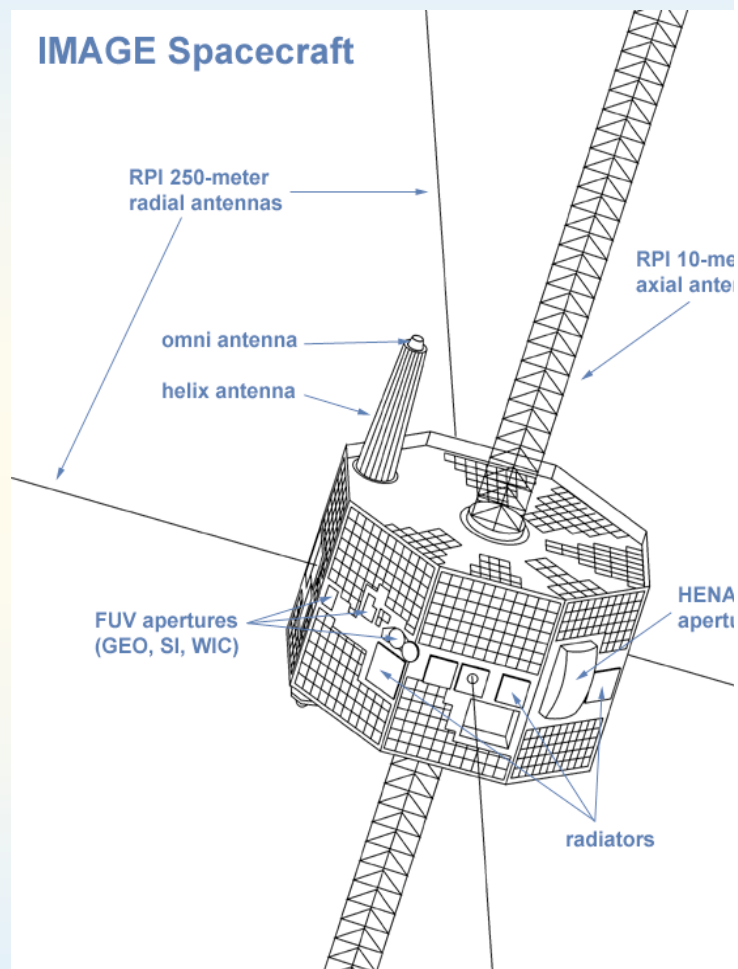
# NASA Satellite Anomaly Resolution



- Imager for Magnetopause-to-Aurora Global Exploration (IMAGE)
- Spacecraft anomaly:
  - No telemetry after 2005-DEC-18
  - No response after revival attempts
  - Does satellite maintain a spin-stabilized attitude?
  - Can spacecraft still receive/execute uplinked commands?
- Objective:
  - Obtain AMOS observations to help diagnose satellite status & anomaly

## State of satellite antennae:

1 intact  
1 missing (lost on orbit)  
2 significantly truncated





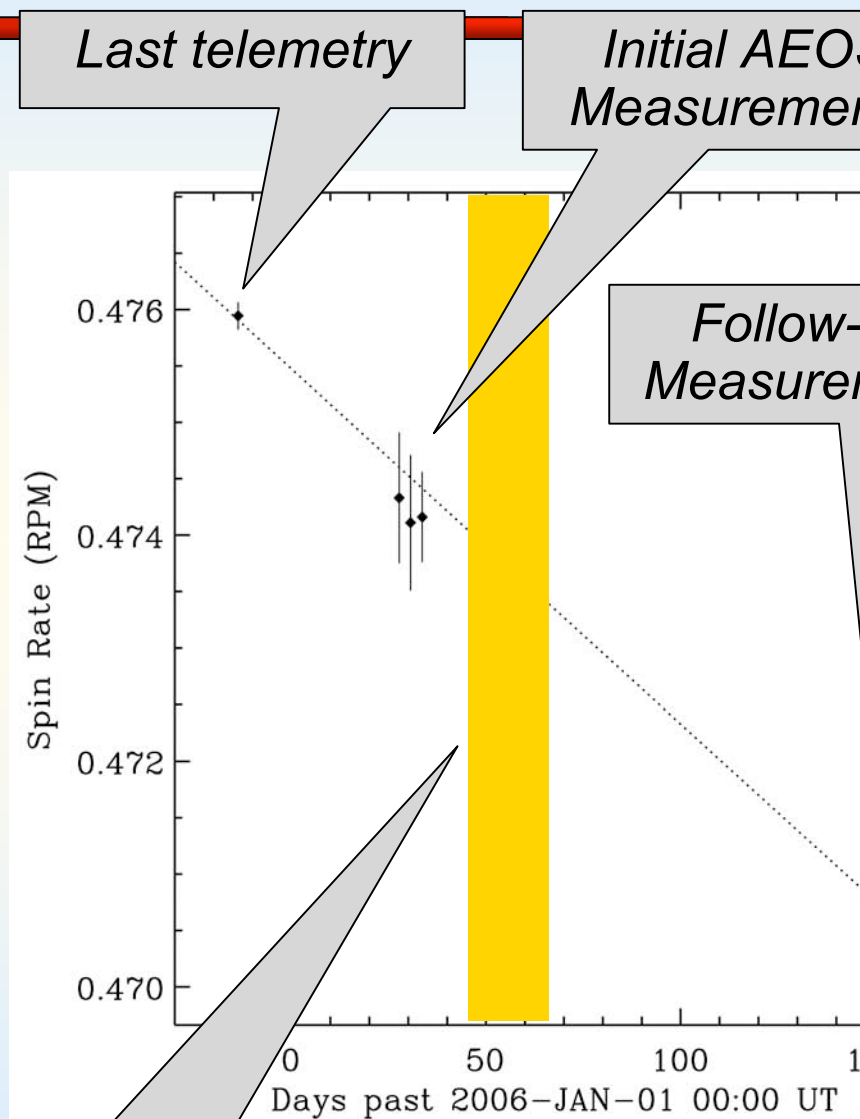
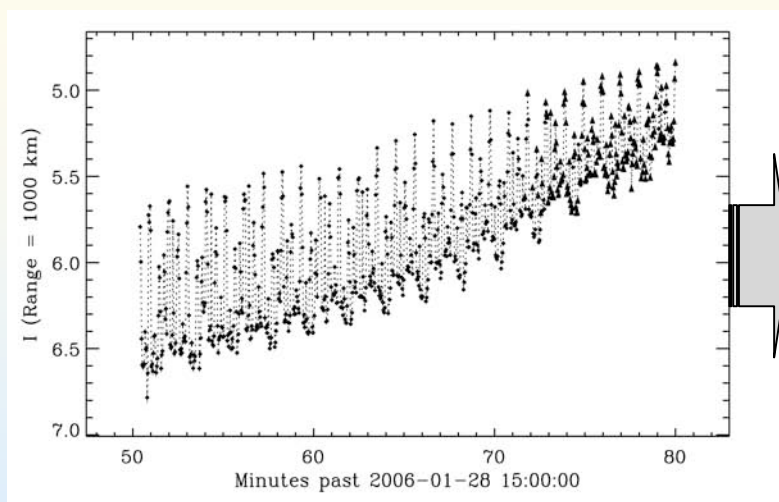
# Confirmed satellite unresponsive



- Can satellite receive and execute commands?

- AMOS measured spin rate
- NASA commanded satellite to spin up
- AMOS measured spin rate

- Conclusion: IMAGE unresponsive to up-linked commands and steadily spinning down



Spin-up commands up-linked (0.52 rpm)

